Flexible Depth Completion for Sparse and Varying Point Densities (Supplementary Material)

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A. Overview

In this supplementary, we provide additional dataset details, quantitative results, and qualitative visualizations. These sections are organized as follows:

- Section B provides additional details regarding the KITTI and NYUv2 depth completion datasets. Further, we provide additional details and visualization of our selected scan-lines on KITTI for our few-line setting.
- Section C contains additional implementation and training details of our Affinity-Based Shift Correction (ASC) module.
- Section D extends our ablation study in the main paper with tables containing additional metrics as well as with visualizations of intermediate depth maps.
- Sections E and F contain a more extended leaderboard comparison on the 64-line KITTI and 500-point NYUv2 settings.
- Section G presents a comparison of our method to Sparsity Agnostic Depth Completion [5].
- Section I contains additional details of the nuScenes depth completion dataset as well as visualizations for this difficult domain adaptation setting.
- Section J contains further discussion regarding our experiments using stronger monocular depth estimation models.
- Section K provides results of applying 3D detection method VoxelRCNN on the completed depth maps and demonstrates a single pipeline for 3D detection on variable sparsity LiDAR.

B. Additional Dataset Details

B.1. KITTI Dataset

The KITTI dataset [5, 12] contains 87k pairs of 64-line Li-DAR depth maps and RGB images for training, 1,000 images for selected validation, and 1,000 images for online testing. We use the 1,000-image selected validation set in our experiments. With few LiDAR points at the top, the images are bottom-cropped to 256x1216 for training and testing similar to prior work. [23, 28, 35, 43, 45].

B.2. Scan Line Selection on KITTI Dataset

As the 64-line LiDAR in the KITTI dataset is quite dense, with each image pixel being within 5 pixels of a depth point, we evenly subsample the 64-line LiDAR to simulate more affordable fewer-line LiDAR sensors. To get scan lines, we transform points to spherical world coordinates and bin by zenith angle similar to prior work. In doing so, we find that depth completion performance varies greatly for 1, 2, 4, and 8 line LiDAR depending on the pitch of the simulated sensor. As the LiDAR sensor is carefully placed to ensure maximum scene coverage in realistic scenarios, we similarly carefully choose the best setup for each 1, 2, 4, and 8 line simulated sensor to maximize depth completion performance. As we are downsampling 64-line LiDAR to simulate fewer-line LiDAR, "pitch" of the simulated sensor is represented by choosing different line indicies (out of the 64 lines in the original KITTI sensor) for the first scan line of the fewer-line sensor. For instance, choosing index 0 for the 1-line LiDAR causes the single scan line to be pointed directly at the immediate ground location, resulting in poor depth completion performance. On the other hand, an index of 63 places the single scan line above most elements of the scene, similarly resulting in poor performance.

To ensure that the selected scan lines are not biased to any evaluated depth completion model, we use a separate model to determine the most suitable scan lines. More specifically, we use a separate model with the same standard architecture as MIDAS [31], a commonly used monocular model, but with RGBD input and a ResNet50 backbone. Note that while this model shares the same architecture as MIDAS, the pre-trained MIDAS weights are not used, and the backbone only has ImageNet pre-training. A model is trained for each # of scan lines over varying pitches.

The results of this model evaluated on various pitches is shown with different metrics in Figure 5. We find that depth completion performance varies greatly over different pitches for few-line sensors. Notably, we find that a poorly placed 4-line sensor (2542 RMSE at starting index 15) can even be outperformed by a optimally placed 2-line sensor (2414 RMSE at starting index 22), demonstrating that the



Figure 5. We show depth completion performance for various chosen pitches of simulated few-line sensors. In practice, the pitch of the simulated sensor corresponds to the line index (out of the 64 lines in the original KITTI LiDAR sensor) of the first scan-line in the simulated few-line LiDAR sensor). We find that depth completion performance varies greatly with this choice for few-line sensors.

selection of pitch for the simulated fewer-line is critical. To closely mirror real-world settings where such few-line sensors are placed carefully, we select the best setup for each few-line sensor, yielding starting indices 53, 22, 9, and 3 for 1, 2, 4, and 8-line LiDAR, respectively. We re-emphasize that our selection of these scan-lines is not biased to any of the models we evaluate - we had used a separate architecture mimicking MIDAS, but with RGBD input, to select scan lines. Visualizations of our optimally selected scan-lines and other sub-optimally selected scan-lines are shown in Figure 6. We will release generated sparse depth maps for training and evaluation, and we hope future work compare on this same, more realistic few-line LiDAR setting.

B.3. NYUv2 Dataset

The NYUv2 dataset [32] contains 120k RGB-D images collected by a Microsoft Kinect sensor in 464 indoor scenes. We follow previous work [2, 27, 28, 35, 43] and train on 50k images from the training set and evaluate on the 654 images from the official test set. Images are downsampled and center cropped to 304x228.

C. Additional Implementation Details

We train our model with a batch size of 24 and a learning rate of 2e-4 on NYUv2, and a batch size of 8 and learning rate of 3.3e-4 on KITTI. We use the AdamW [20, 26] optimizer with weight decay 1e-2. The depth loss weight α is 0.1, 0.15, 0.25, and 0.5 for 16x, 8x, 4x, and 2x resolution decoder stages, and is decayed in the later epochs of



Figure 6. The chosen scan lines generally have better coverage of more diverse and distant scene elements. The depth predictions are from the MIDAS-like RGBD model used to select the scan lines. Error maps using KITTI's error color scheme are visualized below each depth map prediction.

training. The partially scale-invariant loss [9] is used for $\mathcal{D}^{initial}$. For \mathcal{D}^{fuse} and \mathcal{D}^{final} , we use ℓ_1 for NYUv2 and both ℓ_1 and ℓ_2 for KITTI following prior work [28]. Attention using Flash Attention [6, 7] and RoPE [34] is used for both the transformer encoder and cross-attention layers [37] in the ASC module. If the input sparse depth map has more than 5500 points, which is the average # of points for 16-line LiDAR, 5500 points are independently randomly sampled for the ASC modules at each scale.

Regarding weighted sum of point features, we note that SparseFormer [38] fuses a single-channel sparse feature they interpret as "confidence." We find that because this feature does not receive direct supervision, it has little correlation with prediction quality. Thus, we add more channels and interpret it as a deep feature.

D. Extended Ablation Study

D.1. Full Ablation Study Tables

For completeness and to allow future work to fully compare with various settings of our framework, we extend our ablation studies in Tables 1, 2, 3, and 4 from the main paper with full precision and additional metrics in Tables 11, 12, 13, and 14, respectively. Note that in Table 13 we additionally include application of the NLSPN head on our pipeline with RGB input. We find that it improves performance for ResNet34+ backbone but worsens performance for ResNet34 and Effb5. As such, when using RGB-input, we do not apply add the NLSPN head.

D.2. Visualization of Weighted Sum Depth Maps

To further validate the importance of taking a weighted sum over depth errors intead of raw input depths, we visualize intermediate attention maps and and weighted sum depth



Figure 7. Visualization of intermediate attention maps and weighted sum depth predictions for models with weighted sum over features & depths and features & depth errors.

	# of Sampled Points														
Components	2				8	8 32 200					500				
	$\delta_{1.25}\uparrow$	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓
Features	0.8849	0.1060	0.4138	0.9273	0.0764	0.3477	0.9590	0.0500	0.2678	0.9867	0.0238	0.1591	0.9930	0.0168	0.1175
Depths	0.8150	0.1361	0.5042	0.8904	0.0985	0.4056	0.9405	0.0643	0.3091	0.9681	0.0390	0.2116	0.9745	0.0331	0.1800
Depth Errors	0.8380	0.1300	0.4551	0.9146	0.0844	0.3587	0.9611	0.0496	0.2561	0.9877	0.0229	0.1499	0.9935	0.0159	0.1107
Features + Depths	0.8830	0.1075	0.4187	0.9224	0.0783	0.3539	0.9586	0.0504	0.2690	0.9856	0.0251	0.1648	0.9921	0.0183	0.1247
Features + Depth Errors	0.8709	0.1152	0.4279	0.9247	0.0782	0.3476	0.9613	0.0492	0.2612	0.9873	0.0230	0.1534	0.9934	0.0160	0.1128

Table 11. Ablation on different uses of affinity with full precision and metrics.

	# of Sampled Points														
Components	2				8			32			200		500		
	$\delta_{1.25}\uparrow$	REL↓	RMSE↓												
Features + Depth Errors	0.8709	0.1152	0.4279	0.9247	0.0782	0.3476	0.9613	0.0492	0.2612	0.9873	0.0230	0.1534	0.9934	0.0160	0.1128
+ ω_{fuse} w/o \mathcal{F}^{dist}	0.8716	0.1126	0.4273	0.9234	0.0786	0.3493	0.9618	0.0485	0.2592	0.9876	0.0227	0.1523	0.9936	0.0157	0.1108
$+ \tilde{\mathcal{F}}^{dist}$	0.8821	0.1083	0.4131	0.9261	0.0783	0.3452	0.9620	0.0488	0.2580	0.9873	0.0228	0.1524	0.9934	0.0159	0.1118
+ Partial SI-Loss for $\mathcal{D}^{initial}$	0.8790	0.1081	0.4236	0.9287	0.0766	0.3450	0.9644	0.0472	0.2517	0.9879	0.0225	0.1492	0.9937	0.0158	0.1100
+ Initial Feature-only fusion	0.8872	0.1024	0.4094	0.9321	0.0730	0.3348	0.9632	0.0473	0.2533	0.9873	0.0231	0.1530	0.9934	0.0161	0.1121
+ ℓ_1 for \mathcal{D}^{final} and \mathcal{D}^{fuse}	0.8946	0.0980	0.4001	0.9376	0.0688	0.3263	0.9662	0.0440	0.2450	0.9875	0.0220	0.1520	0.9934	0.0153	0.1117

Table 12. Ablation on correction confidence, initial feature-only fusion, and loss functions with full precision and metrics.



Figure 8. Additional visualizations of intermediate attention & depth maps.

Figure 9. Additional visualizations of intermediate attention & depth maps.

			# of Sampled Points														
Backbone	-D Input	NLSPN		2			8			32			200			500	
			$\delta_{1.25}\uparrow$	REL↓	RMSE↓												
Res34	X	X	0.8638	0.1100	0.4434	0.9192	0.0780	0.3605	0.9603	0.0470	0.2606	0.9875	0.0217	0.1512	0.9934	0.0151	0.1112
Res34	✓	X	0.8179	0.1338	0.5078	0.9070	0.0863	0.3814	0.9617	0.0481	0.2588	0.9879	0.0223	0.1481	0.9934	0.0161	0.1113
Res34	X	1	0.8614	0.1112	0.4545	0.9173	0.0795	0.3709	0.9602	0.0479	0.2640	0.9869	0.0228	0.1554	0.9932	0.0162	0.1142
Res34	✓	1	0.8154	0.1318	0.5123	0.9067	0.0842	0.3812	0.9620	0.0467	0.2572	0.9883	0.0208	0.1444	0.9939	0.0141	0.1058
Res34+	X	×	0.8154	0.1309	0.5034	0.9069	0.0856	0.3830	0.9612	0.0490	0.2597	0.9871	0.0232	0.1529	0.9927	0.0168	0.1169
Res34+	✓	×	0.8210	0.1327	0.4957	0.9162	0.0805	0.3576	0.9681	0.0435	0.2369	0.9901	0.0198	0.1350	0.9948	0.0140	0.1006
Res34+	X	1	0.8150	0.1298	0.5055	0.9106	0.0837	0.3780	0.9640	0.0472	0.2541	0.9879	0.0225	0.1497	0.9931	0.0163	0.1142
Res34+	~	1	0.8350	0.1272	0.4749	0.9215	0.0783	0.3473	0.9690	0.0425	0.2335	0.9903	0.0190	0.1325	0.9951	0.0130	0.0970
Original	I NLSPN	Model	0.8220	0.1321	0.4973	0.9117	0.0844	0.3664	0.9668	0.0444	0.2399	0.9899	0.0194	0.1349	0.9949	0.0131	0.0980
Effb5	X	X	0.8946	0.0980	0.4001	0.9376	0.0688	0.3263	0.9662	0.0440	0.2450	0.9875	0.0220	0.1520	0.9934	0.0153	0.1117
Effb5	✓	×	0.8841	0.1013	0.4159	0.9336	0.0706	0.3311	0.9660	0.0445	0.2440	0.9875	0.0229	0.1512	0.9931	0.0168	0.1144
Effb5	X	1	0.8941	0.1000	0.4066	0.9347	0.0707	0.3299	0.9662	0.0447	0.2461	0.9874	0.0228	0.1529	0.9930	0.0165	0.1150
Effb5	1	1	0.8934	0.0977	0.3994	0.9363	0.0687	0.3234	0.9679	0.0425	0.2386	0.9881	0.0208	0.1461	0.9939	0.0140	0.1062

Table 13. Ablation on backbones, inputs, and NLSPN with full precision and metrics.

							# of S	Sampled I	Points						
Components		2			8			32			200			500	
	$\delta_{1.25}$ \uparrow	REL↓	RMSE↓	$\delta_{1.25}$ \uparrow	REL↓	RMSE↓	$\delta_{1.25}$ \uparrow	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓	$\delta_{1.25}\uparrow$	REL↓	RMSE↓
R34+ RGBD	0.8210	0.1327	0.4957	0.9162	0.0805	0.3576	0.9681	0.0435	0.2369	0.9901	0.0198	0.1350	0.9948	0.0140	0.1006
w/ CSPN	0.8297	0.1332	0.4833	0.9204	0.0802	0.3507	0.9696	0.0428	0.2325	0.9904	0.0193	0.1328	0.9950	0.0134	0.0980
w/ NLSPN	0.8350	0.1272	0.4749	0.9215	0.0783	0.3473	0.9690	0.0425	0.2335	0.9903	0.0190	0.1325	0.9951	0.0130	0.0970

Table 14. Ablation on refinement he	ad with full	precision and	metrics.
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predictions for a model taking a weighted sum over features & depths and one taking a weighted sum over features & depth errors. These models correspond to rows 4 and 5 in Table 1, respectively. The visualizations are shown in Figures 7, 8, and 9. To visualize the attention map, for each pixel, we show the color of the depth point it had the highest affinity with. For weighted sum predictions, we visualize the depth map fused back into the decoder features, which is the weighted sum over input depths for the first model shown and the weighted sum over depth *errors* applied to intermediate predictions for the second model shown. Note that for fair comparison with the ablative baseline of summing over input depths, the model summing over errors does not leverage our proposed confidence correction module.

We observe that a weighted sum over depths produces depth maps of low quality, when compared to depth maps of a weighted sum over depth offsets. This can be attributed to two main factors. First, in settings with few # of points, the input points often cannot cover the entire range of depths in the scene. For example, in Figure 7 in the 2 and 8 point settings, the depths of the back of the library are not captures in the sparse input depth. As such, simply interpolating between input depth values to generate intermediate predictions is insufficient to produce reasonable depth maps as they cannot reach beyond the boundaries of the closest and furthest depths. Instead taking a weighted sum of depth errors - differences between the predicted and input depths - and applying those corrections to intermediate predictions can cover the entire range of depths in the scene.

Second, however, we find that even when the # of points increases, covering the full range of depths in the scene, the weighted sum over input depths still has errant depth values where a weighted sum over depth errors does not. More specifically, considering the 200 and 500 point settings in Figure 7, we see some artefacts near the top of the scene for weighted sum of depth predictions. This is primarily caused by errors in the cross attention. When some pixels attend to the wrong input points, potentially due to incorrect semantic groupings, taking a weighted sum over input depths directly transfers those irrelevant points' depths to those pixels, resulting in completely incorrect depth predictions. However, instead taking a weighted sum over depth errors and applying a correction to intermediate predictions, depth predictions for those pixels are still largely grounded by those pixels' intermediate depth predictions, only slightly offset potentially incorrectly by irrelevant corrections. Such small errors can more easily be corrected later on in the model when semantics become more clear, while large errors as those caused by weighted sum over input depths are more readily propagated incorrectly to final predictions. This is especially clear in settings with more # of points where the primary source of errors is from predictions at depth boundaries, where such mistakes are most common. As such, we find that taking a weighted sum of depth offsets consistently outperforms taking a weighted sum of depths es-

Method	RMSE	MAE	iRMSE	iMAE
CSPN [3]	1019.6	279.4	2.93	1.15
Sparse&Dense [18]	917.6	234.8	2.17	0.95
BDBF [30]	900.3	216.4	2.37	0.89
TWISE [17]	840.2	195.5	2.08	0.82
NConv [10]	829.9	233.2	2.60	1.03
S2D [27]	814.7	249.9	2.80	1.21
FusionNet [36]	772.8	215.0	2.19	0.93
DepthNormal [41]	777.1	235.2	2.42	1.13
DSPN [42]	766.7	220.3	2.47	1.03
MSG-CHN [22]	762.2	220.4	2.30	0.98
DeepLiDAR [29]	758.3	226.5	2.56	1.15
FuseNet [2]	752.9	221.2	2.34	1.14
ACMNet [45]	744.9	206.0	2.08	0.90
CSPN++ [4]	743.6	209.2	2.07	0.90
PointFusion [15]	741.9	201.1	1.97	0.85
NLSPN [28]	741.6	199.5	1.99	0.84
ENet [14]	741.3	216.3	2.14	0.95
GuideNet [35]	736.2	218.8	2.25	0.99
FCFRNet [24]	735.8	217.1	2.20	0.98
PENet [14]	730.0	210.5	2.17	0.94
RigNet [43]	712.6	203.2	2.08	0.90
DySPN [23]	709.1	192.7	1.88	0.82
CompFormer [44]	708.9	203.5	2.01	0.88
Ours	727.3	194.3	1.96	0.83

Table 15. Online Test Set Evaluation for 64-line KITTI.

pecially when there are more input points.

E. Extended Comparison on 64-line KITTI

In the main paper, we primarily focused on the fewer-line and variable sparsity settings for KITTI. While not our focus, we also provide online test set results for comparison. Results are in Table 15. We observe that although our ASC module is primarily developed for sparse and variable point settings, our pipeline can achieve competitive results in the 64-line online test set. Furthermore, we have shown that our module can be applied to any encoder-decoder model and is complementary to advancements in depth completion as shown through our experiments with various spatial propagation heads in the main paper.

F. Extended Comparison on 500-point NYUv2

Similarly, in addition to our extensive experiments on few and variable point settings in the main paper, we provide a comparison with existing work on the 500-point setting for NYUv2. Results are in Table 16. Our pipeline performs competitively with existing work on this largely saturated benchmark. We emphasize that in more difficult settings with fewer points and variable input distributions, our

Method	$\delta_{1.25}\uparrow$	REL↓	RMSE↓
CSPN [3]	0.992	0.016	0.117
CSPN++ [4]	-	-	0.116
DeepLiDAR [29]	0.993	0.022	0.115
ACMNet [45]	0.994	0.015	0.105
Plane-Residual [21]	0.994	0.014	0.104
SparseFormer [38]	0.994	0.014	0.104
DepthCoeff [16]	0.994	0.013	0.118
DepthNormal [41]	0.995	0.018	0.112
GNN [40]	0.995	0.016	0.106
FCFRNet [24]	0.995	0.015	0.106
GuideNet [35]	0.995	0.015	0.101
PointFusion [15]	0.996	0.014	0.090
CostDCNet [19]	0.995	0.013	0.096
TWISE [17]	0.996	0.013	0.097
RigNet [43]	0.996	0.013	0.090
NLSPN [28]	0.996 (0.9955)	0.012 (0.0117)	0.092 (0.0924)
DySPN [23]	0.996	0.012	0.090
GraphCSPN [25]	0.996	0.012	0.090
CompFormer [44]	0.996	0.012	0.090
Ours	0.996 (0.9956)	0.012 (0.0115)	0.092 (0.0917)

Table 16. 500-point setting evaluation on NYUv2.

method far outperforms state-of-the-art as discussed in the main paper. Such settings are important for wider applicability of depth completion models for other datasets and tasks, and we encourage future work to evaluate on such sparser and variable distribution settings as well.

G. Comparison to Sparsity Agnostic Depth Completion

We evaluate on the sparse depth maps generated and released by SpAgNet [5]. Results are shown in Table 17. First, when trained on 500 points, our proposed framework with RGB-inpu outperforms SpAgNet is most settings. Unlike SpAgNet which does global scale correction regardless of input point location or semantic information, our pipeline considers each point individually and does semantics-guided shift corrections. Completely agnostic to the location of each point, SpAgNet is more robust to extreme distributional differences from the 500 points seen during training, and it transfers slightly better to shifted grid and 5 point settings. However, by catering to each point, our pipeline demonstrates better performance in all other settings.

Then evaluating various methods trained on 2 to 500 randomly sampled points, we find that performance improves for all settings, most notably even for the uneven shifted grid and Livox patterns, which have very different distributions compared to random sampling. This corroborates our findings from evaluating on SIFT keypoint distributions

	Mathad	Shifte	d Grid	Livox	Pattern	5 P	oints	50 F	oints	100	Points	200	Points	oints 500 Points	
	Method	REL↓	RMSE↓	REL↓	RMSE↓	$\text{REL} \downarrow$	RMSE↓	REL↓	RMSE↓	REL↓	RMSE↓	$\text{REL} \downarrow$	RMSE↓	REL↓	RMSE↓
ğ	pNCNN [11]	0.519	1.922	0.061	0.333	0.722	2.412	0.108	0.568	0.061	0.338	0.040	0.237	0.026	0.170
Ξ.	CSPN [3]	0.367	1.547	0.066	0.376	0.581	2.063	0.185	0.884	0.067	0.388	0.027	0.177	0.016	0.118
Τü	NLSPN [28]	0.175	0.796	0.037	0.233	0.262	1.033	0.081	0.423	0.038	0.246	0.019	0.142	0.013	<u>0.101</u>
ats	PackNet-SAN [13]	-	-	-	-	-	-	-	-	-	-	0.027	0.155	0.019	0.120
ö	SpAgNet [5]	0.110	0.422	0.039	0.206	0.131	0.467	0.058	0.272	0.038	0.209	0.024	0.155	0.015	0.114
0 F	Ours (NLSPN Base)	0.190	0.832	0.046	0.264	0.262	0.892	0.097	0.435	0.046	0.269	0.020	0.140	0.013	0.096
50	Ours (R34 RGB)	<u>0.131</u>	<u>0.539</u>	0.030	0.186	<u>0.145</u>	<u>0.584</u>	0.044	0.247	0.030	0.191	0.022	0.149	<u>0.015</u>	0.110
2	NLSPN [28]	0.080	0.356	0.026	0.162	0.102	0.423	0.036	0.209	0.026	0.168	0.020	0.134	0.014	0.101
50	Ours (NLSPN Base)	0.078	0.344	0.025	0.158	0.095	0.398	0.035	0.202	0.026	0.164	0.019	0.132	0.013	0.100
6	Ours (R34 RGB)	0.070	0.332	0.029	0.181	0.090	0.398	0.039	0.228	0.029	0.187	0.022	0.152	0.016	0.115

Table 17. Evaluation on sparse depth maps from SpAgNet [5]. Bottom three rows are trained on $2 \sim 500$ points.

Mathad	1 L	ine	4 Li	nes	16 L	ines	64 Lines		
Method	RMSE	MAE↓	RMSE	MAE↓	RMSE	MAE↓	RMSE	MAE↓	
NLSPN	3507.7	1849.1	2293.1	831.3	1288.9	377.2	889.4	238.8	
DySPN	3625.5	1924.7	2285.8	843.3	1274.8	366.4	878.5	228.6	
CompletionFormer	3250.2	1582.6	2150.0	740.1	1218.6	337.4	848.7	215.9	
Ours (NLSPN Base)	3039.6	1365.7	2116.5	678.5	1206.7	324.2	818.2	205.3	

Table 18. Eval on various # of scan-lines on KITTI. Metric is mm.

that models trained on randomly sampled 2 to 500 points can transfer well even to unique patterns and distributions of input points.

H. Comparison to CompletionFormer

In this section, we additionally compare with CompletionFormer [44]. For fair comparison, we re-train on the sub-splits and sparse depth maps released by Completion-Former. Our results are in Table 18, with baseline results taken from Table 4 of CompFormer. We find that our pipeline outperforms prior work, especially significantly for the sparsest 1-line sensor. This shows the importance of our ASC module's flexible pixel-point interaction.

I. Transfer Performance on nuScenes

In the main paper, we demonstrated quantitatively in Table 9 that 1) models with an RGBD-input encoder trained on just 64-lines on KITTI perform poorly when transferred to 32-line nuScenes [1] depth completion. 2) Our pipeline, applied to a simple ResNet34 backbone encoder-decoder with an RGB-input encoder transfers much better under similar training settings. 3) Training on variable 1 to 64 lines on KITTI yields more robust models that perform much better when transferred not only to the base 32-line nuScenes dataset but also to the simulated 8-line and 16-line nuScenes LiDAR. 4) Our pipeline with the ASC module outperforms NLSPN with or without an RGB-input encoder.

We verify these findings qualitatively in Figures 10, 11, and 12. We first notice that RGBD-input encoder models, NLSPN and Ours (NLSPN Base), trained on 64-line KITTI generate artifacts in their depth maps when applied to nuScenes as noted by prior work [39]. These artefacts follow the input distribution, indicating that 64-line trained RGBD models are not able to handle the increased distance between image pixels and depth points caused by fewer-line LiDAR and higher resolution images in nuScenes. We do note, however, that our module applied to NLSPN significantly reduces the extent of these artifacts and improves performance (MAE \downarrow). Furthermore, we find that our ASC module applied to a standard R34 RGB-input encoder architecture transfers very well, far outperforming both RGBD models, generating largely coherent structures, and not producing any line artefacts for 32-line LiDAR. We do see performance degrade for 16 and 8 lines as the domain shift increases.

We then verify that training on variable lines on KITTI yields robust, transferable models by training all three methods on 1 to 64 lines on KITTI. We find that performance improves for all settings and that the generated depth maps are of higher quality without line artefacts. Notably, we observe that our RGB-input encoder pipeline still performs the best, demonstrating that our ASC module is able to adaptively propagate depth information even under significant domain shift. Additionally, these experiments suggest that fusing sparse depth information at the input layer, as is common in prior works, may result in worse performance when transferring to different domains, compared to using an RGB-input encoder and fusing depth later. We hope that our proposed ASC module serves as a strong baseline for further investigations in this direction.

J. Scaling Mono. Depth Estimation Models

In Tables 12 and 13 in the main paper, we show that our pipeline is complementary to larger backbones and large-scale MiDaS [31] mono-depth pretraining for both depth completion and joint estimation and completion. Notably, completion performance for sparser regimes (2 and 32 points) increases steadily as we apply the ASC module to stronger pre-trained backbones, showing that our module can effectively align these strong, context-based monocular predictions with sparse point input. On the other hand, in the dense 500 point setting where most pixels are within



Figure 10. Domain Adaptation from KITTI to nuScenes. "64-Line" represents models trained with 64 lines on KITTI, and "1 to 64-Line" indicates the model was trained on variable sparsity, randomly sampling from 1 to 64 lines on KITTI. We emphasize that nuScenes data was not seen by any model during training. Error maps using KITTI's error color scheme are visualized below each depth map prediction.



Figure 11. Additional Visualizations of Domain Adaptation from KITTI to nuScenes.



Figure 12. Additional Visualizations of Domain Adaptation from KITTI to nuScenes.

a few pixels of some input point, a simple, high-res backbone with RGBD input (ResNet34+ with RGBD) outperforms stronger backbones with RGB input (MiDaS BeiT-L). Based on our ablations in Table 3 in the main paper, this is because processing depth with the CNN encoder and maintaining higher resolution (removing initial 4x downsampling) is crucial for this dense setting.

K. Single Pipeline for 3D Detection on Variable Scan Lines

We use our proposed depth completion model to generate dense depth maps and outproject them to a 3D point cloud. We then concatenate the original LiDAR points with the depth completed point cloud, adding a channel for a flag indicating whether the point is from the LiDAR sensor or the depth completion model. For our 3D detector, we adopt VoxelRCNN [8] for its strong performance and efficiency. Note that we took care to remove from the depth completion training set sequences geographically close to samples in the 3D detection validation set as mentioned by [33].

The results are presented in Table 19. We show mAP at moderate difficulty and 0.7 IoU threshold for the most common Car class. Echoing our analyses in the main paper, we find that our depth completion model consistently improves performance over just LiDAR at all sparsity levels. We note that in the very few-line settings of 1 or 2 lines, the LiDARonly model largely collapses, unable to make reasonable predictions. Our completion-then-detection pipeline can detect some cars even in this setting, mainly close cars most immediately relevant for autonomous driving. Furthermore, this completion-then-detection pipeline largely maintains performance even when using a single pipeline for variable scan-lines. Finally, in the extreme case where a single pipeline is trained for 64 lines and deployed to fewer scan-lines, we find that the completion-then-detection pipeline stays far more robust than the LiDAR-only detection pipeline. We hypothesize that inputting the densified point cloud into the model leads to a far smaller domain shift in terms of point density and number between 64-line and fewer-line settings compared to just using the raw Li-DAR point clouds. In all, we demonstrate that our depth completion significantly improves downstream 3D detection and can be effectively leveraged for a completion-thendetection pipeline over variable sparsities.

Training Setup	Depth Completion	1 Line	2 Lines	4 Lines	8 Lines	16 Lines	32 Lines	64 Lines
Each # of scan lines	X	0.10	1.89	28.32	49.63	66.11	77.79	84.03
Each # of scan lines	1	11.99	26.22	48.17	63.28	75.31	81.73	84.37
Variable # of scan lines	X	0.03	1.49	22.91	47.74	65.46	77.63	81.19
Variable # of scan lines	1	10.03	25.82	49.40	60.47	73.16	79.01	81.73
Only 64 lines	X	-	-	0.51	19.89	49.90	72.24	84.03
Only 64 lines	1	10.32	17.49	40.37	53.53	69.84	79.07	84.37

Table 19. 3D detection performance using the proposed depth completion model on KITTI.

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