# Lighting Every Darkness in Two Pairs: A Calibration-Free Pipeline for RAW Denoising

—— Supplementary Materials -

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https://srameo.github.io/projects/led-iccv23

## **A. Network Architecture**

We illustrate the detailed network architecture of the proposed LED in Fig. 1, a UNet-style [6] architecture with five stages. Both in encoder and decoder, each stage is consisted of two sequential RepNR blocks. It is worth noting that, except AINDNet [3], all other methods shared a same UNet architecture as SID [1]. Moreover, the LED would finally yield the same architecture for fair comparison after reparameterization (Sec. B).

## **B. Structural Reparameterization Process**

In this section, we would detail the process of structural reparameterization. As stated in Sec. 3.4 in the main paper, the RepNR block consists of serial vs. parallel linear mapping, which can be fused to a single one. Specifically, the RepNR block can be transformed into a plain  $3 \times 3$  convolution. Formally, this architecture contains two  $3 \times 3$  convolutions with weights  $\{W_0, W_1\}$  and bias  $\{b_0, b_1\}$ , and one of them follows a CSA layer. Let CSA(x) = kx + b, where k, b denote the weight and bias of it. Thus, the result for the input x can be represented as:

$$\begin{split} \tilde{x} &= W_0(CSA(x)) + b_0 + W_1 x + b_1 \\ &= W_0(kx+b) + b_0 + W_1 x + b_1 \\ &= (W_0 k + W_1) x + (W_0 b + b_0 + b_1) \\ &= \tilde{W}x + \tilde{b}, \end{split}$$
(1)

where the whole deployment process is formulated. It demonstrates that our RepNR block can be transformed into a plain  $3 \times 3$  convolution, and brings no extra costs during inference. It worth noting that we leveraged the online reparameterization strategy same as [2], thus there is no performance gap at all between training and testing.

## C. More Visual Results

LED could better recover details compared with ELD [7] (calibration-based method) in Fig. 2. As shown in Fig. 3, LED outperforms other calibration-based methods [7, 8] in removing out-of-model noise. In Fig. 4 and Fig. 5-8, we provide more results on two benchmarks: ELD [7] and SID [1]. The restoration results of Kristina *et al.* [5], Noise2Noise [4], AINDNet [3], Zhang *et al.* [8] and ELD [7] are presented for comparison.

## References

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Figure 1. Detailed network architecture for our proposed LED. The  $\hat{C} \times \hat{H} \times \hat{W}$  formatted expression on the arrow indicates the feature size for the corresponding stage.  $H \times W$  is the input resolution. RepNR block with  $\times 2$  denotes two RepNR blocks in a sequential way. After structural reparameterization (Sec. B), our method outputs a same structure as SID [1] and other methods for fair comparison.



(a) Input

## (b) ELD [7]

(c) LED (Ours)





Figure 3. Compared with state-of-the-art calibration-based methods: ELD [7] and Zhang *et al.* [8], proposed LED is able to remove the out-of-model noise (*Zoom-in for best view*).



Figure 4. Visual comparison between our LED and other state-of-the-art methods on the ELD [7] dataset (*Zoom-in for best view*). We amplified and post-processed the input images with the same ISP as ELD [7].









Input / PSNR

Kristina et al. [5] / 36.78

Noise2Noise [4] / 40.37

AINDNet [3] / 40.31



Zhang et al. [8] / 40.33



ELD [7] / 40.71



LED (Ours) / 41.32



GT /  $\infty$ 



Input / PSNR



Kristina et al. [5] / 40.09



Noise2Noise [4] / 41.91



AINDNet [3] / 41.57



Zhang et al. [8] / 41.46





ELD [7] / 41.27



LED (Ours) / 42.86









Zhang et al. [8] / 43.31





LED (Ours) / 44.33

GT /  $\infty$ 

Figure 5. Visual comparison between our LED and other state-of-the-art methods on the SID [1] dataset (Zoom-in for best view). We amplified and post-processed the input images with the same ISP as ELD [7].







ELD [7] / 39.28



AINDNet [3] / 39.70



GT /  $\infty$ 



Input / PSNR



Kristina et al. [5] / 42.47



LED (Ours) / 42.14

Noise2Noise [4] / 46.61



AINDNet [3] / 44.30



Zhang et al. [8] / 46.69

Zhang et al. [8] / 42.67



ELD [7] / 47.68

ELD [7] / 39.72



LED (Ours) / 47.69

LED (Ours) / 43.56



GT /  $\infty$ 

 Input / PSNR
 Kristina et al. [5] / 39.96

 Noise2Noise [4] / 42.78
 AINDNet [3] / 41.86

Figure 6. Visual comparison between our LED and other state-of-the-art methods on the SID [1] dataset (*Zoom-in for best view*). We amplified and post-processed the input images with the same ISP as ELD [7].



Input / PSNR



Kristina et al. [5] / 42.87



Noise2Noise [4] / 45.32



AINDNet [3] / 44.38



Zhang et al. [8] / 45.10



ELD [7] / 45.26



LED (Ours) / 46.44



GT /  $\infty$ 



Input / PSNR



Kristina et al. [5] / 35.65



Noise2Noise [4] / 41.64



AINDNet [3] / 38.93



Zhang et al. [8] / 40.93





LED (Ours) / 40.93



GT /  $\infty$ 



Input / PSNR



Kristina et al. [5] / 44.18



Noise2Noise [4] / 45.41





Figure 7. Visual comparison between our LED and other state-of-the-art methods on the SID [1] dataset (Zoom-in for best view). We amplified and post-processed the input images with the same ISP as ELD [7].



Input / PSNR



Zhang et al. [8] / 37.25



Kristina et al. [5] / 35.46

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LED (Ours) / 37.53



AINDNet [3] / 37.09



GT /  $\infty$ 



Input / PSNR



ELD [7] / 37.29

Kristina et al. [5] / 43.10



Noise2Noise [4] / 44.72



AINDNet [3] / 44.70



Zhang et al. [8] / 46.54



Input / PSNR





Kristina et al. [5] / 32.97



LED (Ours) / 48.77

Noise2Noise [4] / 38.80



GT /  $\infty$ 



AINDNet [3] / 38.40



Figure 8. Visual comparison between our LED and other state-of-the-art methods on the SID [1] dataset (Zoom-in for best view). We amplified and post-processed the input images with the same ISP as ELD [7].