# Frequency-Aware Self-Supervised Monocular Depth Estimation Supplementary Material

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### 1. Pseudo-code of Ambiguity-Masking

The overall algorithm of the proposed Ambiguity-Masking is summarized as Alg. 1. Please refer to our Github repository for full implementation.

Algorithm 1 Extract Ambiguities for Photometric Loss

- **Input:** target image  $I_t$ , source images  $I_{t+n}$ , indices of source images  $src_ids$ , reconstructed images  $\tilde{I}_{t+n}$ , photometric errors of all source images  $\mathcal{L}$
- **Output:**  $A_t^{pe}$ : ambiguity mask of the final photometric error
- 1:  $A_t \leftarrow \text{EXTRACTAMBIGUITYFORIMAGE}(I_t);$
- 2:  $reproj\_ambiguities \leftarrow list;$
- 3: for all n in  $src_i ds$  do
- 4:  $\mathcal{A}_{t+n} \leftarrow \text{EXTRACTAMBIGUITYFORIMAGE}(I_{t+n});$
- 5:  $\tilde{\mathcal{A}}_{t+n} \leftarrow$  bilinear sample  $\mathcal{A}_{t+n}$  subject to  $\otimes_{t+n}$ ; // to get which pixels in reconstructed  $\tilde{I}_{t+n}$  are from the ambiguous pixels in source  $I_{t+n}$ .
- 6: append  $\tilde{\mathcal{A}}_{t+n}$  to reproj\_ambiguities;
- 7: end for
- 8: min\_idx ← argmin(L); // we adopt min. reprojection loss from [12].
- A'<sub>t</sub> ← reproj\_ambiguities[min\_idx]; // to gather ambiguity value adopted in the final loss map.
- 10:  $\mathcal{A}_t^{max} \leftarrow \max(\mathcal{A}_t, \mathcal{A}'_t); // \text{ as Eq. 13.}$

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11: \mathcal{A}_t^{pe} \leftarrow \mathcal{A}_t^{max} < \delta; // as Eq. 14.
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- 12: return  $\mathcal{A}_{t}^{pe}$ ;
- 13: **procedure** EXTRACTAMBIGUITYFORIMAGE(*I*)
- 14:  $\mathcal{F} \leftarrow \text{compute frequency map of } I; // \text{ as Eq. 9.}$ 15:  $\mu \leftarrow \nabla_{u+} \cdot \nabla_{u-} < 0 \mid | \nabla_{v+} \cdot \nabla_{v-} < 0; // \text{ as Eq. 10.}$
- 16:  $\mathcal{A} \leftarrow \mu \mathcal{F};$
- 17: return  $\mathcal{A}$ ;
- 18: end procedure

#### 2. Further Consideration on the Two Modules

We let the Ambiguity-Masking module take input from the Auto-Blur because we want the high-freq regions of input images to be first processed by Auto-Blur before extracting ambiguities. The reason for this lies in the fact that without smoothing the high-frequency areas, the Ambiguity-Masking would wrongly filter out almost all pixels in high-frequency areas as the *dense thin* objects inside are likely to be misjudged as ambiguous colors, disabling them from participate in training.

#### 3. Full Numbers of Hyper-params Ablation

In this section, we show full numbers of ablations of all hyper-parameters in our methods, as reported in Tab. 1. We then give detailed analyses on each one of them.

If  $\delta$  is too small, the Amb.-masking will wrongly exclude some non-ambiguous pixels, *e.g.*, the long wall from near to far could also satisfy the constraint of gradual color transition, but it does not belong to the problem demonstrated in Fig. 1. If  $\delta$  is too large, boundaries with little color difference will be missed.

For kernel size s in Auto-Blur, if we decrease s, the receptive field could not be effectively enlarged when measuring pixel similarity. If we increase s too much, the central pixel's contribution (its own characteristic color) is reduced since the Gaussian distribution gets 'shorter' and 'wider'.

For threshold  $\lambda$ , decreasing  $\lambda$  would wrongly smooth the texture-less regions, as the already-weak supervision signal on them will be further weakened. Increasing  $\lambda$  too much would miss some pixels in high-freq regions which could confuse the photometric loss as illustrated in Fig. 2.

For the percentage threshold  $\eta$  of high-frequency pixels in Auto-Blur, when  $\eta$  is too small, not only the textureless regions but also some object boundary areas which does not belong to 'high-frequency area' would be wrongly smoothed. When  $\eta$  is too large, the same as  $\lambda$ , our Auto-

Hyper-parameter	Value	Abs Rel	Sq Rel	RMSE	RMSE log	$\delta {<} 1.25$	$\delta < 1.25^2$	$\delta < 1.25^3$
δ	$\begin{array}{c c} 0.2 \\ 0.3 \\ 0.4 \end{array}$	0.113 0.112 0.113	0.884 <b>0.834</b> 0.864	4.814 <b>4.746</b> 4.757	0.190 <b>0.189</b> 0.190	0.878 <b>0.880</b> 0.879	0.960 <b>0.961</b> 0.960	0.982 0.982 0.982
s	7 9 11	0.112 0.112 0.113	0.836 <b>0.834</b> 0.868	4.753 <b>4.746</b> 4.782	0.190 <b>0.189</b> <b>0.189</b>	0.878 <b>0.880</b> 0.877	<b>0.961</b> <b>0.961</b> 0.960	0.981 0.982 0.982
λ	$\begin{array}{c c} 0.15 \\ 0.20 \\ 0.25 \end{array}$	0.113 0.112 0.113	0.844 <b>0.834</b> 0.881	4.814 <b>4.746</b> 4.797	0.192 <b>0.189</b> 0.191	0.879 <b>0.880</b> 0.877	0.959 <b>0.961</b> 0.959	0.982 0.982 0.981
η	50 60 70	0.113 0.112 0.114	0.860 <b>0.834</b> 0.887	4.804 <b>4.746</b> 4.839	0.192 <b>0.189</b> 0.190	0.875 <b>0.880</b> 0.878	0.959 <b>0.961</b> 0.960	0.981 0.982 0.982

Table 1. Ablations on all hyper-parameters.



Figure 1. High-resolution qualitative comparisons of Monodepth2 [12] with and w/o our proposed methods (input from CityScapes [3]).

Blur would be too strict, *i.e.* miss to smooth some pixels in high-frequency areas which could confuse the photometric loss.

## 4. Full-Resolution Qualitative Results

We show more full-resolution qualitative depth predictions in Fig. 1 (CityScapes) and Fig. 2 (KITTI).



Figure 2. High-resolution qualitative comparisons of *Depth-Hints* [32] and *Monodepth2* [12] with and w/o our proposed methods (input from KITTI [10]).