Supplementary Material: Radiometric Calibration from Faces in Images

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1. Pseudo-albedo \hat{A} for Single Pigment Density Change Identification

A key step to our approach is identifying the single pigment density change map C and we describe it in Sec. 5.1 in the main paper. We use the pseudo-albedo \hat{A} by applying intrinsic image decomposition on the uncalibrated image Iinstead of the skin albedo A. We show that the an accurate identification of C can be obtained only using \hat{A} .

Considering the nearby pixel pair (p,q) satisfying the single pigment density change assumption, thus

$$A_q = \sigma^{\Delta} A_p, \tag{1}$$

where σ is the relative absorbance vectors of the changing pigment and Δ is the density difference. Further assume the shading of (p,q) are same, so by applying the camera response function f, we have

$$I_p = f(A_p * D), \qquad (2)$$

$$I_q = f(\sigma^{\Delta} A_p * D), \qquad (3)$$

where D is the diffuse shading of (p,q).

Applying the intrinsic image decomposition on uncalibrated image I, we can decompose I_p and I_q into the pseudo-albedo \hat{A} and pseudo-shading \hat{D}

$$I_p = \hat{A}_p * \hat{D}, \tag{4}$$

$$I_q = \hat{A}_q * \hat{D}. \tag{5}$$

Combing Eq. (2), Eq. (3), Eq. (4), and Eq. (5), we have

$$\hat{A}_p * \hat{D} = f(A_p * D), \tag{6}$$

$$\hat{A}_q * \hat{D} = f(\sigma^{\Delta} A_p * D).$$
(7)

Taking the ration of Eq. (6) and Eq. (7), we have

$$\frac{\hat{A}_p * \hat{D}}{\hat{A}_q * \hat{D}} = \frac{f(A_p * D)}{f(\sigma^{\Delta} A_p * D)}.$$
(8)

Let considering a certain class, gamma function $f(x) = x^{\gamma}$, of the camera response function f, then

$$f(\sigma^{\Delta}A_p * D) = (\sigma^{\Delta}A_p * D)^{\gamma}, \tag{9}$$

$$= \sigma^{\Delta \cdot \gamma} (A_p * D)^{\gamma}, \qquad (10)$$

$$= \sigma^{\Delta \cdot \gamma} f(A_p * D). \tag{11}$$

Combing Eq. (11) and Eq. (8) and eliminated the same factors, \hat{D} and $f(A_p * D)$, appearing in both numerators and denominators, we have

$$\frac{\hat{A}_p}{\hat{A}_q} = \frac{1}{\sigma^{\Delta \cdot \gamma}}.$$
(13)

Denoting $\hat{\Delta} = \Delta \cdot \gamma$, we have

$$\hat{A}_q = \sigma^{\hat{\Delta}} \hat{A}_p. \tag{14}$$

Comparing Eq. (14) with Eq. (1), the single pigment density change property of pixel pair (p, q) in skin albedo A is preserved in its pseudo-albedo \hat{A} . As a result, the single pigment density change map C can be identified only using the pseudo-albedo \hat{A} , without known skin albedo A. Although the aforemention formula derivation only for a ceratin type of camera response function, our experiment results show C can be identified from \hat{A} for any other camera response functions, not limited to a ceratin gamma functions.

2. Real Image Results

We present more real image results in this supplementary material. A quantitative error comparison over different cameras, namely DSLRs (*Canon EOS 5D*, *Nikon D5100*), MILCs (*SONY NEX-7*, *SONY a6000*), webcams (*Logitech Pro9000*, *ANC 152WS*), and mobile phone cameras (*iPhone 6 Plus*, *HUAWEI Mate8*), is given in Table 1. The results for each individual camera is shown from Fig. 1 to Fig. 8.

The results of EdgeCRF [2], GICRF [3] and RankCRF [1] depend on image content and their ability to find suitable local regions for their processing. Especially in outdoor images full of complex surfaces, there may be a number of erroneously detected regions

^{*}This work was done while Chen Li was an intern at Microsoft Research.

Cameras	EdgeCRF [2]	GICRF [3]	ISCRF [4]	RankCRF [1]	Ours w.o. shading	Ours
Canon EOS 5D	11.3	16.7	5.91	4.47	5.5.61	3.29
Nikon D5100	20.5	15.9	11.2	11.0	5.69	4.03
SONY NEX-7	26.8	27.5	13.4	13.0	7.90	3.22
SONY α 6000	30.9	28.5	14.5	23.2	5.35	3.37
Logitech Pro9000	30.9	26.6	4.62	17.0	5.80	4.05
ANC 152WS	26.3	17.5	5.34	10.8	7.30	3.45
iPhone 6 Plus	16.0	22.0	5.19	4.11	4.70	2.71
HUAWEI Mate8	33.3	32.7	21.2	22.0	4.90	3.85

Table 1. Average radiometric calibration error for real images ($\times 10^2$).

that do not actually satisfy the assumed properties and thus mislead the estimation. Although RankCRF [1] rely on the same scene radiance property as EdgeCRF [2], it performs a bit better than EdgeCRF [2] because the rank minimization is more robust at avoiding trivial solutions. GICRF [3] utilizes the locally planar region whose radiance have be approximated linearly. However, the number of these locally planar region is limited due to the nonuniform illumination, reflections, and complex textures in real world. ISCRF [4] works more effectively for cameras that generate more image noise, such as a webcam or a smartphone camera, and generally less well for higher quality images. In this experiment, ISCRF [4] happens to work well for some of the Canon EOS 5D images because of the blurred backgrounds due to depth-of-field and a higher ISO used for capturing indoor images with low illumination.

When a face is present, our approach can obtain relatively reliable estimates regardless of the other scene content, even if the face regions are not large, are captured under varying illumination, or are noisy. This is indicated by the average error statistics in Tab. 1. From these examples of typical images, it can be seen that faces are a useful source of radiometric calibration information that is effectively exploited by our technique.

References

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Figure 1. Real image results for *Canon EOS 5D*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 2. Real image results for *Nikon D5100*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 3. Real image results for *SONY NEX-7*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 4. Real image results for SONY $\alpha 6000$. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 5. Real image results for *Logitech Pro9000*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 6. Real image results for *ANC 152WS*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 7. Real image results for *iPhone 6 Plus*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.



Figure 8. Real image results for *HUAWEI Mate8*. Each raw shows a non-linear input image on the left, followed by the estimated inverse response functions for various methods as well as the ground truth for the R, G and B channels, respectively. Please zoom in for better viewing.